

SQL Queries

Modern astronomical surveys

- With the advent of the digital age, driven by the use of CCDs in cameras, astronomical surveys have started to become semi or fully automated
 - So, huge amounts of data are now arriving from sky surveys (Tb=Terabyte, Pb = Petabyte = 1000Tb)
 - ~50Tb of reduced data products over ~10 years of the Sloan Digital Sky Survey (SDSS)
 - ~2Pb of reduced data products over ~5 years of the Dark Energy Survey (DES)
 - ~50Pb of reduced data products over ~5 years of LSST (Large Synoptic Survey Telescope) operations
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Mining modern astronomical surveys

- With such a large amount of data to sift through, astronomers have become more involved in developing data mining techniques
 - We've discussed aspects of this in terms of pixelating the sky...which is really a method for indexing large amounts of data in a database for efficient searches
 - The HTM index, a type of quad-tree that we've discussed briefly, is an efficient schema for storing data and searching through that data by object position
 - We won't discuss the math of HTM in detail (a good description is linked from the syllabus) but think of it as a HEALPix-like index, coupled with the spherical cap formalism to find which HTM pixels lie in a cap
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Introduction to SQL

- Visit the *SDSS SQL Tutorial* linked from the syllabus
 - Read and/or try the following tutorials:
 - 1. *Introduction*
 - 2. *A simple Query*
 - 3. *Common Searches*
 - Note though, that nothing in these first 3 SQL tutorials makes use of the HTM indexing scheme
 - The genius of HTM is coded in functions such as, e.g., *fGetNearbyObjEq(α, δ, θ)* which can *very* rapidly find objects at a radius θ around a position (α =RA, δ =dec)
 - Try *SDSS SQL Tutorial 10. Functions*
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Python tasks

1. Using the *SDSS SQL Search Box* (linked from the syllabus) download RA, Dec and g-band imaging for all objects in the SDSS within $\theta = 2'$ of $(\alpha, \delta) = (300^\circ, -1^\circ)$
 2. Write Python code that reads in these SDSS objects and plots RA against dec. *Use circles for your data points*
 3. Repeat your plot, but bin your points such that objects with *larger* g are plotted using *smaller* circles (i.e. plot $16 < g < 17$ at one size, plot $17 < g < 18$ at smaller size)
 - Using Matplotlib, `plt.scatter(ra, dec, s=s)` will allow you to plot points of different sizes
 - Here, *s* is the “size” of the point, but note that it’s actually the *area* of the marker (so multiplying *s* by 4 will double the radius of the plotted point)
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Python tasks

4. Use the *SDSS Navigator Tool* linked from the syllabus, to display the SDSS image near $(\alpha, \delta) = (300^\circ, -1^\circ)$. Zoom in until the image is a few arcminutes across (the scale will be about $20''$)
 - Check that your plot looks reasonably like the *SDSS Navigator Tool* image