

# SQL Queries

# Modern astronomical surveys

---

- With the advent of the digital age, driven by the use of CCDs in cameras, astronomical surveys have started to become semi or fully automated
  - So, huge amounts of data are now arriving from sky surveys (Tb=Terabyte, Pb = Petabyte = 1000Tb)
    - ~50Tb of reduced data products over ~10 years of the Sloan Digital Sky Survey (SDSS)
    - ~2Pb of reduced data products over ~5 years of the Dark Energy Survey (DES)
    - ~50Pb of reduced data products over ~5 years of LSST (Large Synoptic Survey Telescope) operations
-

# Mining modern astronomical surveys

---

- With such a large amount of data to sift through, astronomers have become more involved in developing data mining techniques
  - We've discussed aspects of this in terms of pixelating the sky...which is really a method for indexing large amounts of data in a database for efficient searches
  - The HTM index, a type of quad-tree that we've discussed briefly, is an efficient schema for storing data and searching through that data by object position
  - We won't discuss the math of HTM in detail (a good description is linked from the syllabus) but think of it as a HEALPix-like index, coupled with the spherical cap formalism to find which HTM pixels lie in a cap
-

# Introduction to SQL

---

- Visit the *SDSS SQL Tutorial* linked from the syllabus
  - Read and/or try the following tutorials:
    - 1. *Introduction*
    - 2. *A simple Query*
    - 3. *Common Searches*
  - Note though, that nothing in these first 3 SQL tutorials makes use of the HTM indexing scheme
  - The genius of HTM is coded in functions such as, e.g., *fGetNearbyObjEq( $\alpha, \delta, \theta$ )* which can *very* rapidly find objects at a radius  $\theta$  around a position ( $\alpha$ =RA,  $\delta$ =dec)
    - Try *SDSS SQL Tutorial 10. Functions*
-

# Python tasks

---

1. Using the *SDSS SQL Search Box* (linked from the syllabus) download RA, Dec and g-band imaging for all objects in the SDSS within  $\theta = 2'$  of  $(\alpha, \delta) = (300^\circ, -1^\circ)$
  2. Write Python code that reads in these SDSS objects and plots RA against dec. *Use circles for your data points*
  3. Repeat your plot, but bin your points such that objects with *larger* g are plotted using *smaller* circles (i.e. plot  $16 < g < 17$  at one size, plot  $17 < g < 18$  at smaller size)
    - Using Matplotlib, `plt.scatter(ra, dec, s=s)` will allow you to plot points of different sizes
    - Here, *s* is the “size” of the point, but note that it’s actually the *area* of the marker (so multiplying *s* by 4 will double the radius of the plotted point)
-

## Python tasks

---

4. Use the *SDSS Navigator Tool* linked from the syllabus, to display the SDSS image near  $(\alpha, \delta) = (300^\circ, -1^\circ)$ . Zoom in until the image is a few arcminutes across (the scale will be about  $20''$ )
  - Check that your plot looks reasonably like the *SDSS Navigator Tool* image